

1) Notes 58(8): Simple Example of the Effect of Torsion on Deflection Angle

The angle of deflection is given by:

$$\delta = 2 \left(\frac{v_{rr} - 1}{r} \right)_{r=r_0} \quad (1)$$

In the absence of torsion:

$$d \Lambda v_{rr,0} = -\omega_0 \Lambda v_{rr,0} \quad (2)$$

where ω_0 is the spin connection in the absence of torsion.

In the presence of torsion:

$$d \Lambda v_{rr,T} = -\omega \Lambda v_{rr,T} + \delta T \quad (3)$$

If δT is considered to be a small perturbation then it is a first approximation:

$$\omega \sim \omega_0 \quad (4)$$

and

$$\omega \Lambda v_{rr,T} \sim \omega_0 \Lambda v_{rr,0} \quad (5)$$

$$\text{so } d \Lambda v_{rr,T} - d \Lambda v_{rr,0} \sim \delta T \quad (6)$$

$$\propto \boxed{d \Lambda (\delta v_{rr}) \sim \delta T} \quad (7)$$

$$\text{and } \boxed{\Delta \delta \sim 2 \left(\delta v_{rr} \right)_{r=r_0}} \quad (8)$$

2) From eqs. (7) and (8) it is clear that the torsional perturbation δT will change the angle of deflection by $\Delta \delta$. In Cartesian coordinates introduce a perturbation:

$$\delta q_{Vrr} = \frac{1}{\sqrt{2}} (1-i) e^{i\phi} \quad - (9)$$

$$\sim \frac{1}{\sqrt{2}} \sin \phi \sim \frac{1}{\sqrt{2}} \phi \quad - (10)$$

for $\phi \ll 1$. Thus:

$$\Delta \delta \sim \frac{2}{\sqrt{2}} \phi \quad - (11)$$

This is a very simple result that illustrates the effect of torsion on the angle of deflection in an Edington experiment. From experimental data on gravitational lensing it is known that ϕ must be very small for the sun-planet system because the EH result is accurate here to one part in 10^5 . For other systems ϕ may be larger.

In this very simple example the δq_{Vrr} is due to a helical spacetime perturbation.