

324(7) : Velocity Curve of a Whirlpool Galaxy  
 The orbit of a star in a whirlpool galaxy is  
 the hyperbolic spiral:

$$\frac{1}{r} = \frac{\theta}{r_0} \quad - (1)$$

Its velocity from eq. (2) of Note 324(6) is :

$$v^2 = \frac{L^2}{m^2} \left( \left( \frac{d}{d\theta} \left( \frac{1}{r} \right) \right)^2 + \frac{1}{r^2} \right) \quad - (2)$$

$$\frac{1 + \frac{L^2}{m^2 c^2} \left( \left( \frac{d}{d\theta} \left( \frac{1}{r} \right) \right)^2 + \frac{1}{r^2} \right)}{1 + \frac{L^2}{m^2 c^2} \left( \left( \frac{d}{d\theta} \left( \frac{1}{r} \right) \right)^2 + \frac{1}{r^2} \right)}$$

$$= \frac{L^2}{m^2} \left( \frac{1}{r_0^2} + \frac{1}{r^2} \right) \quad - (3)$$

$$1 + \frac{L^2}{m^2 c^2} \left( \frac{1}{r_0^2} + \frac{1}{r^2} \right)$$

Therefore :

$$v \xrightarrow{r \rightarrow \infty} \frac{L}{mr_0} \left( 1 + \frac{L^2}{m^2 c^2 r_0^2} \right)^{-1/2} \quad - (4)$$

= constant.

The theory gives the correct  $r \rightarrow \infty$  of the  
 velocity curve of a whirlpool galaxy.

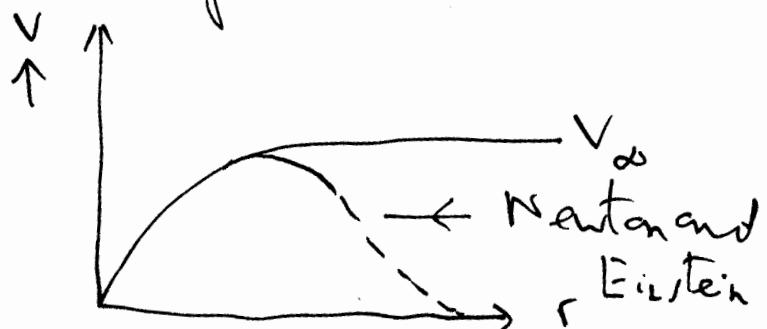
By definition:

$$v^2 = r^2 + \dot{\theta}^2 r^2 - (5)$$

So  $v \rightarrow 0$  — (6)

The velocity curve is therefore sketched in the following:

Fig.(1)



where

$$v_\infty = \frac{L}{mr_0} \left( 1 + \frac{L^2}{r_0^2 c^2 r_0^2} \right)^{-1/2} - (7)$$

The Newtonian velocity is well known to be:

$$\begin{aligned} v^2 (\text{Newton}) &= mg \left( \frac{2}{r} - \frac{1}{a} \right) \\ &= mg \left( \frac{2}{r} - \frac{1}{a} (1 - e^{-2}) \right) \\ &= \frac{mg}{r} \left( 2 + \frac{(e^2 - 1)}{1 + e \cos \theta} \right) - (8) \end{aligned}$$

So  $v(\text{Newton}) \rightarrow 0$  — (9)

as sketched in Fig (1).

The Einsteinian orbit is:

$$r = \frac{d}{1 + e \cos(\chi \theta)} - (10)$$

which is a precessing ellipse or conic section, so in eq. (8) :

$$v(\text{Einstein}) = \frac{mg}{r} \left( 2 + \frac{(e^2 - 1)}{1 + e \cos(\varphi\theta)} \right) - (11)$$

$\xrightarrow[r \rightarrow \infty]{ } 0$

So S.O.Q. Newton and Einstein fail completely  
in whirlpool galaxies where ECE2 describes  
the velocity curve straightforwardly from eqs. (2)  
and (5), given the observed orbit (1).

In the non-relativistic limit the integral  
for of Brie's equation is eq. (5) of note 324(4) :

$$U = H - \frac{L^2}{2m} \left( \left( \frac{d}{d\theta} \left( \frac{1}{r} \right) \right)^2 + \frac{1}{r^2} \right) - (12)$$

$$\text{so } U = H - \frac{L^2}{2m} \left( \frac{1}{r_0^2} + \frac{1}{r^2} \right) - (13)$$

for a hyperbolic spiral orbit. The force is

$$F = - \frac{dU}{dr} = - \frac{L^2}{mr^3} - (14)$$

This result is also given by the Brie's equation,

$$F = -\frac{L^2}{mr^3} \left( \left( \frac{d^2}{d\theta^2} \left( \frac{1}{r} \right) + \frac{1}{r} \right) - (15) \right)$$

Because from eq. (1)

$$\frac{d^2}{d\theta^2} \left( \frac{1}{r} \right) = 0 - (16)$$

Therefore the potential is:

$$U = \int \frac{L^2}{mr^3} dr - (17)$$

$$= -\frac{L^2}{2mr^2}$$

From eqs (13) and (17):

$$-\frac{L^2}{2mr^2} = H - \frac{L^2}{2mr_0^2} - \frac{L^2}{2mr^2} - (18)$$

so

$$H = \boxed{\frac{L^2}{2mr_0^2}} - (19)$$

This is the classical hamiltonian of the hyperbolic spiral orbit of a star in a spiral galaxy.

### 3) Conclusions

For the orbit of a star in a whirlpool galaxy:

$$\frac{1}{r} = \frac{\theta}{r_0} - (20)$$

$$F = -\frac{L^2}{mr^3} - (21)$$

$$U = -\frac{L^2}{2mr^2} - (22)$$

$$H = \frac{L^2}{2mr_0^2} - (23)$$

$$v_\infty = \frac{L}{mr_0} \left( 1 + \frac{L^2}{m^2 c^2 r_0^2} \right)^{-1/2} - (24)$$

Resonates spiral orbits. The star is attracted inward toward the centre of the galaxy and falls in to the centre in a spiral defined by eq. (20). There is no black hole at the centre of the galaxy, and there is no dark matter. The velocity of the star at infinite distance from the centre is given by eq. (24) and is constant.