

241(4): Definition of the Factor x from the Michowski Equation

In the limit of a nearly circular orbit the force of the Michowski equation is:

$$\underline{F} = \left(-\gamma^4 \frac{m \underline{M} G}{r^2} - \gamma^2 \frac{d m \underline{M} G (1 - \gamma^2)}{r^3} \right) \underline{e}_r \quad (1)$$

For the true precessing ellipse:

$$r = \frac{d}{1 + \epsilon \cos(x\theta)} \quad (2)$$

$$\underline{F} = \left(-\frac{m \underline{M} G x^2}{r^2} - d (1 - x^2) \frac{m \underline{M} G}{r^3} \right) \underline{e}_r \quad (3)$$

(Comparing eqs. (1) and (3):

$$x^2 = \frac{\gamma^4 + d \gamma^2 (1 - \gamma^2) - d}{r - d} \quad (4)$$

Let $\gamma^2 = \left(1 - \frac{v^2}{c^2} \right)^{-1} \quad (5)$

and $v^2 = \underline{M} G \left(\frac{2}{r} - \frac{1}{a} \right), \quad (6)$

$$a = \frac{d}{\epsilon^2 - 1} \quad (7)$$

Now calculate t as a function of r as follows.

First, note that:

$$2) \quad L_0 = m r^2 \frac{d\theta}{dt} = m r^2 \frac{d\theta}{dr} \frac{dr}{dt} \quad - (8)$$

From eq. (2): $\frac{dr}{d\theta} = \frac{\epsilon x r^2 \sin(x\theta)}{d} \quad - (9)$

and $\cos(x\theta) = \frac{1}{\epsilon} \left(\frac{d}{r} - 1 \right) \quad - (10)$

with $\cos^2(x\theta) + \sin^2(x\theta) = 1 \quad - (11)$

So $\sin(x\theta) = \left(1 - \frac{1}{\epsilon^2} \left(\frac{d}{r} - 1 \right)^2 \right)^{1/2} \quad - (12)$

and $\frac{dr}{d\theta} = \frac{\epsilon x r^2}{d} \left(1 - \frac{1}{\epsilon^2} \left(\frac{d}{r} - 1 \right)^2 \right)^{1/2} \quad - (13)$

and $\frac{d\theta}{dr} = \frac{d}{\epsilon x r^2} \left(1 - \frac{1}{\epsilon^2} \left(\frac{d}{r} - 1 \right)^2 \right)^{-1/2} \quad - (14)$

From eqs. (8) and (14):

$$L_0 = \frac{m d}{\epsilon x} \left(1 - \frac{1}{\epsilon^2} \left(\frac{d}{r} - 1 \right)^2 \right)^{-1/2} \frac{dr}{dt} \quad - (15)$$

3) So:

$$dt = \frac{md}{cL_0} \left(1 - \frac{1}{c^2} \left(\frac{d}{r} - 1 \right)^2 \right)^{-1/2} dr - (16)$$

and

$$t = \frac{md}{cL_0} \int \frac{1}{x} \left(1 - \frac{1}{c^2} \left(\frac{d}{r} - 1 \right)^2 \right)^{-1/2} dr - (17)$$

in which x is given by eqs. (4) and (5).

In the limit of a nearly circular orbit:

$$a \sim r - (18)$$

and

$$v^2 \sim \frac{MG}{r} - (19)$$

so

$$\gamma^2 = \left(1 - \frac{MG}{c^2 r} \right)^{-1} - (20)$$

Computer algebra is needed to integrate eq. (7) and find r in terms of t .
